# CERUSSITE FROM EVEN-GRAINED RAPAKIVI, SALTVIK, ÅLAND ISLANDS, SW FINLAND

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### Introduction

During mineral separation for age determination of even-grained rapakivi granites, some rare accessory minerals were identified. In this connection fluocerite was identified from the Fjälskär granite stock in Houtskär, SW Finland (cf. Suominen 1988; Lahti and Suominen 1988). Cerussite, PbCO<sub>3</sub>, not previously described from Finland, was found in the even-grained rapakivi granite of Åsbacka, Saltvik, Åland Islands, SW Finland. This sample (age determination sample A763 of the Geological Survey of Finland) is derived from the Ödkarby intrusion in Saltvik (see Bergman 1978). Frosterus and Sederholm (1890) called the intensely red and medium-grained rock Haga granite. The modal analysis and chemical composition of the rock type have been given by Bergman (1981 and 1986).

### Heavy accessory minerals

The heavy (d>3.3) minerals of the Åsbacka

sample were studied under a stereo microscope and identified mainly with XRD-methods. Cerussite occurs as small ( $\emptyset < 0.1 \text{ mm}$ ) anhedral to short-prismatic grains varying in colour from colourless to white and reddish yellow. Six grains of cerussite were recognized under the stereo microscope and hand-picked for further studies.

This even-grained rapakivi granite contains less zircon than the surrounding wiborgitic types of rapakivi granite. According to Nurmi and Haapala (1986), this is a common feature of evengrained rapakivi granites. The zircon crystals of the heavy fraction are colourless or yellow to brown, euhedral, tetragonal prisms.

Other minerals identified from the heavy fraction are fluocerite, black and green anatase, yellow xenotime, chalcopyrite (occasionally altered to bornite and malachite), loellingite, arsenopyrite, molybdenite, pyrite, bastnaesite and rutile.

In connection with the heavy mineral studies, two different types of fluorite were identified: one is transparent colourless to light violet with octahedral crystal habit and the other is yellowish brown and anhedral.



Fig. 1. Energy dispersive spectrum of cerussite from Åsbacka, Saltvik.

# Analytical and X-ray diffraction data on cerussite

The energy dispersive spectrum of the Asbacka cerussite shows only lead with traces of strontium (Fig. 1). The content of Sr is estimated to be about 1 weight %.

X-ray powder diffraction data obtained with the Debye-Scherrer method (camera diameter 11.46 cm) are presented in Table 1. Ni-filtered Cu-radiation was used and silicon was mixed in as an internal standard. Due to broadened reflections and possible doublets the error of measurement has to be noted. Some of the hkl-indices of reflections with small d-values differ from those given in the JCPDS file card no. 5-417 (JCPDS; International Centre for Diffraction Data, 1986). The systematic extinctions were set according to the space group Pmcn given in the above mentioned JCPDS-card (cf. group Pnma of Henry and Lonsdale 1952, p. 151). Only reflections with the hkl-indices coinciding with those given in the JCPDS card were used for unit cell calculations with the least squares method. The error in the last decimal was estimated using the relationships given by Kelsey (1964).

The unit cell dimensions seem generally to be slightly smaller than the values reported for synthetic cerussite (cf. e.g. JCPDS Mineral Powder Diffraction File 1986, p. 198: a = 5.183, b = 8.503 and c = 6.144 Å, V = 270.8 Å<sup>3</sup>).

Table 1. X-ray powder diffraction data on cerussite from Åsbacka, Saltvik.

Intensity <sup>1</sup>	d obs, Å	d calc,	Å hk1	
ms	4.407	4.420	110*	
W	4.243	4.245	020*	
VS	3.590	3.587	111*	
S	3.488	3.491	021*	
S	3.053	3.069	002*	
ms	2.579	2.588	200*	
ms	2.523	2.521	112*	
VS	2.485	2.483	130*	
w	2.209	2.210	220*	
VS	2.081	2.079	221*	
ms	2.005	2.006	041*	
ms	1.972	1.979	202*	
S	1.929	1.930	132*	
S	1.853	1.856	113*	
w	1.794	1.793	222*	
vw	1.745	1.746	042*	
ms	1.630	1,630	311*	
ms	1.586	1.586	241*	
ms	1.560	1.560	151*	
vw	1.536	1.534	004*	
w	1.504	1.504	302	
ms	1.472	1.473	330	
w	1.450	1 449	114	
ms	1.330	1 328	332*	
ms	1 308	1 308	341	
W	1.285	1.285	062	
vw	1.267	1 267	153*	
w	1.244	1.243	044	
W	1.213	1 214	421*	
w	1 186	1 187	351	
	+ sev	eral lines	551	
Cell dimensions		Space	Systematic	
		group	extinctions	
$a = 5.177 \pm 0.004 \text{ Å}$		Pmcn	hk1: no	
$b = 8.490 \pm 0.009 \text{ Å}$			hk0: $h + k = 2n$	
$c = 6.137 \pm 0.004 \text{ Å}$			0k1: no	
$V = 269.7 \text{ Å}^3$			h01: 1 = 2n	

 $^{1}$  vs = very strong

s = strong

ms = medium strong

w = weak

vw = very weak

 $^{2} * =$  used for unit cell calculations

### Isotope composition of cerussite

The isotope composition of lead from three tiny cerussite crystals (unwashed) was analysed and found to be:  ${}^{206}Pb/{}^{204}Pb = 21.407$ ,  ${}^{207}Pb/{}^{204}Pb$ = 16.005 and  ${}^{208}Pb/{}^{204}Pb = 40.302$ .

## **Conclusions and discussion**

The slightly reduced unit cell observed is thought to be due to substitution of Sr for Pb in the cerussite lattice. Because the average interatomic distances of the Sr-O bonds are shorter than the corresponding Pb-O bonds in the metal-oxygen polyhedra, shrinkage of the unit cell is to be expected. According to Speer (1983), the interatomic average distances in the orthorhombic carbonates are 2.636 Å for Sr-O and 2.69 Å for Pb-O. Speer (1983) attributes the differences in the lattice parameters of the orthorhombic carbonates to differences in the ionic radii of the divalent cations. However, the lattice parameters observed for the cerussite from Åsbacka are very close to the values that Sahl (1974) determined for synthetic cerussite. Speer (1983) reports the average C-O distance in strontianite to be 1.285 Å, and in cerussite 1.27 Å. Thus the longer C-O distances in the CO<sub>3</sub> groups of strontianite compared to the corresponding distances in cerussite seem to compensate, at least partly, for the smaller Sr ion.

Chang and Brice (1972) studied the subsolidus phase relations of the aragonite-type carbonates and predicted a complete series of solid solution in the system  $SrCO_3$ —PbCO<sub>3</sub>, at least in the temperature range 400—750°C. On the basis of the graphs of Chang and Brice (1972), the a- and c-axes of the Åsbacka cerussite indicate a Sr content of 10 and 2 mole %, respectively. The b-axis observed falls outside the range given, suggesting some uncertainty in the use of the method for determining the Pb/Sr ratio of cerussite. An immiscibility gap in the solid solution at low temperatures would explain the discrepancies observed.

Cerussite is regarded as a secondary mineral formed during the alteration of galena or, less commonly, other lead minerals (Speer 1983). Robinson (1974) found that cerussite from Tui Mine, New Zealand, formed as a supergene mineral at a temperature of about 30°C. The lead isotope composition of the Åsbacka cerussite shows that a considerable amount of radiogenic lead is incorporated in the sample. Such a composition is typical of secondary minerals and proves the supergene nature of the mineral. No further estimations of the temperature of formation were made, but the very low temperature reported by Robinson (1974) from Tui Mine is thought to apply to the cerussite from Åsbacka as well.

The two types of fluorite crystals observed reflect crystallization at different stages. According to Drugman (1932), octahedral, lightcoloured fluorite is typical of granites and pegmatites. Fluorite of veins occurs as more deeply coloured cubic crystals. The latter type seems to be typical of low-temperature fluorite crystallization.

Fluocerite, now identified from two different even-grained rapakivi granites, seems to be a characteristic accessory mineral of even-grained rapakivi granites as late intrusion phases of plutons and stocks.

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